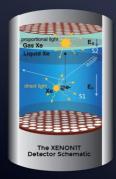
DETECTION PRINCIPLE __

A possible WIMP interaction in the xenon liquid target generates a prompt flash of scintillation light which is recorded by 248 photomultipliers arranged in two arrays: one is mounted in the upper section of the TPC where the gaseous xenon is located, the other one in the lower section, in the liquid xenon.

The active dark matter target, defined by a cylinder with 96 cm height and diameter, is delimited by highly-reflective teflon

Wire electrodes close the cylindrical target from above and below and are used to establish an electric field which prevents the ionization electrons to recombine with parents nuclei and drifts them towards the upper end of the detector.

Massive copper electrodes surrounding the cylindrical target ensure that the electric field lines are perfectly straight. Thus allowing electrons to move through the liquid xenon at a speed of about 2 mm/us.



Once they arrive at the interface between the liquid and the gaseous xenon, a suited phase.

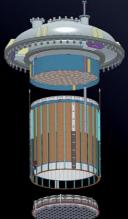
By scattering off the xenon atoms they produce a scintillation signal which is proportional to the number of electrons. These particular type of detectors, which exploit the two phases of the xenon (liquid and gaseous), are referred to as dual-phase time projection chamber.

This light signal is once more recorded by the photomultiplier arrays. The combination between the primary and secondary light signals allows a three dimensional reconstruction of the interaction vertex as well as the selection of events occurring in the central core of the target. The background events, produced by the detector's components, once measured in the outer part of the detector can be excluded, thus reducing the intrinsic background at minimal level.

The ratio of the two signals is further used to distinguish background events. electric field extracts them due to photons and electrons, from a from the liquid into the gas possible dark matter signal interacting with the xenon nuclei.







XENON COLLABORATION: 20 INSTITUTIONS AND MORE THAN 120 PHYSICISTS















Astrophysical observations provide ample evidence for the existence of an invisible even though dominant mass component in the observable universe, from the scales of galaxies up to the largest cosmological structures.

After many decades of increasingly precise astrophysical observations, we have unequivocal evidence that the majority of the material that forms galaxies, clusters of galaxies and the largest observed structures in the cosmos, is non-luminous, or dark (it does not emit or absorb light). Even the precise measurements of the cosmic microwave background radiation and of the abundance of light elements in the Universe, lead to the same result.

According to the current cosmological model, dark matter makes 85% of the total mass of the Universe. The residual mass is the ordinary matter, which we are made of.

Uncovering the nature of this dark matter component is one of the most outstanding challenges in physics today.

The dark matter could be made of a new, yet undiscovered particle, such as a weakly interacting massive particle (WIMP). One way to observe WIMPs is to search for their collisions with atomic nuclei in ultra low-background detectors placed deep underground.

Discovering WIMPs would enlighten two of the outstanding problems of modern physics - the matter composition of our Universe and the existence of new physics beyond the Standard Model. We infer that a few millions of WIMPs are passing through a human body every second, without any interaction.

Nonetheless we can search for WIMPs rare interactions with a variety of experiments on Earth, which are sensitive to the tiny energy deposited in collisions of WIMPs with atoms in a low backaround detector.

The XENONIT experiment, at the Gran Sasso Underground Laboratory, uses liquid xenon to search for these elusive collisions.

Xe XENONIT

The XENONIT experiment is a low-background time projection chamber (TPC), filled with ultra-pure xenon as active detection medium to search for dark matter interactions. The xenon is used in both liquid and gaseous forms. This detector technology offers particularly attractive prospects for dark matter detection due to its excellent capability of rejecting events that are not induced by dark matter interactions and which are therefore called "background" events, as well as its scalability to large mass detectors.

Liquid xenon is an efficient scintillator, i.e., it emits light if a particle hits a xenon atom, and it can also be ionized easily. This allows the simultaneous measurement of a scintillation light and an ionization charge signal for every interaction. Liquid xenon time projection chambers have demonstrated remarkable success in recent years and are currently leading the race towards the direct detection of dark matter.

The xenon is intrinsically radio-pure, stops radiation very efficiently due to its high density, nearly 3 times the one of water, and is sensitive to low-energy interactions.

The XENON project, initiated at the Columbia University, follows a phased approach, with detector masses ranging from tens to hundreds of kilograms of liquid xenon.

In the current phase, the XENONIT detector uses a total of about 3300 kg of ultrapure xenon, out of which 2000 kg are instrumented with 248 novel photomultiplier tubes (light ш detectors) as sensitive sensors for the xenon scintillation 0 light. This inner part of the instrument, the dark matter S target, is surrounded by the remaining xenon, acting as Z radiation shield. In order to have an ultra-low background detector, all materials and components were selected according to the lowest possible intrinsic radioactivity. xenon is a noble element that has gaseous form at room temper, pressure, and liquefies at -108° C. It is present in Earth's atmos, h a very tiny concentration, 0.0000087 % by volume. The xenon's tes it very expensive - its cost is presently more than 1000 \$ per k The liquid xenon is enclosed in a cryostat, which is made of two concentric vessels manufactured with specially selected stainless steel with a very low intrinsic radioactivity. The empty area between the two concentric vessels is necessary to thermally insulate the inner vessel from the environment. The cryostat also accommodates the active detector, the time projection chamber whose signals are recorded by photomultiplier tubes arranged in two arrays, installed above and below the sensitive xenon target. Schematic drawing of Hall B of the LNGS, where XENONIT is located. Details of the "cryostat" are also visible

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The required measurements were conducted with instruments installed in the LNCS laboratory and at other institutions involved in the experiment. These devices have world-leading sensitivities.

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In order to further improve the natural screening capability of the 1400 m of rock overburden provided by the Gran Sasso massif against cosmic ray muons, the cryostat is installed inside a cylindrical tank with 9.6 m diameter and 10 m height where 84 photomultipliers are installed. The tank contains 700 m³ of de-ionized water, that reduces the ambient radioactivity to irrelevant levels. A charged particle passing through the water is able to produce a flash light (Cherenkov radiation) which is detected by the photomultipliers.

Aside the XENONIT experiment a three levels structures is located, hosting all ancillary devices necessary to the functioning of the experiment. The cryogenic system, liquefying the xenon gas, is installed on the top floor, as well as the purification system, which continually recirculates the xenon through special filters. This system removes impurities from the liquid xenon at a level of one atom of impurity per ten billion of xenon atoms.

The counting room, computing and control infrastructure are located at the middle floor of the building.

The bottom floor hosts the custom storage system of the precious stock of xenon. This device can store up 7500 kg of xenon in safe conditions in state of matter - gas, liquid and even solid - without the need for electrical power. In case of an emergency situation, it allows to recover the xenon from the cryostat in few hours. At the same floor there is a 5m high distillation column, which removes from the xenon minute amounts of krypton (one atom per ten thousand billions of Xe atoms) able to produce unexpected events.

All these components, most notably the storage system, are designed such that a detector upgrade to a considerably larger mass is possible relatively quick and easy. Only the inner cryostat vessel and the time projection chamber have to be re-built.



The XENONIT experimental facilities underground in Hall B.